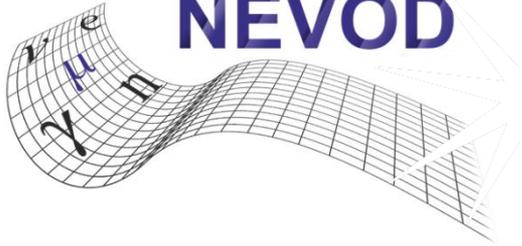


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# **Muonography of the Earth's atmosphere and near-terrestrial space**

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# Introduction

**Muonography** is the method (technique) of investigations of various objects by means of the muon flux.

Analogies: *radiography, electronography, neutronography, protonography*, etc.

Two main fields of investigations by means of *muonography*:

- Heliosphere and magnetosphere of the Earth due to close **connection of muon and primary cosmic ray trajectories**;
- Earth's atmosphere and various objects on the Earth due to a high **penetrating ability of muons**.

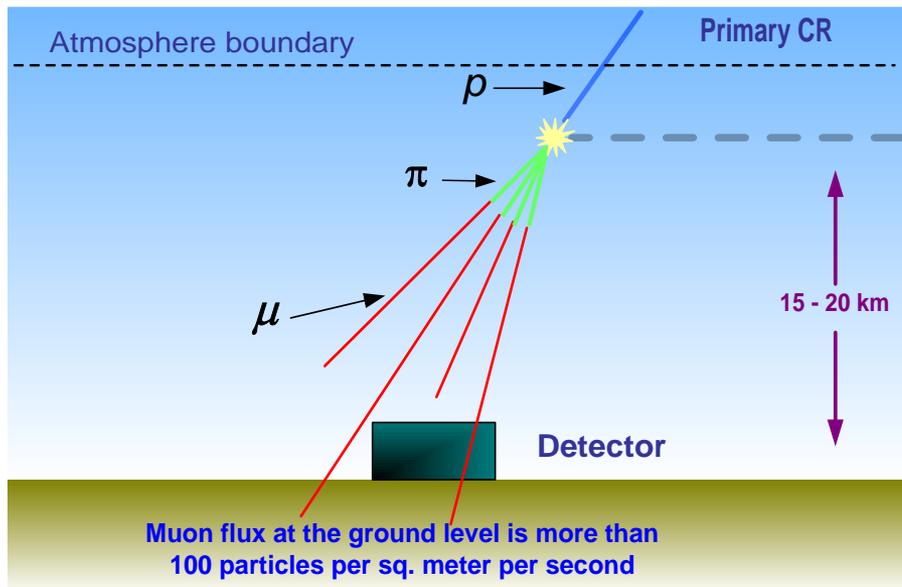
The goal is to give a basic idea of the muonography and its applications.

## Content

- Cosmic ray muons
- From muon detector to muon hodoscopes
- Muonography of heliosphere and magnetosphere
- Muonography of atmosphere and Earth's objects
- Conclusion

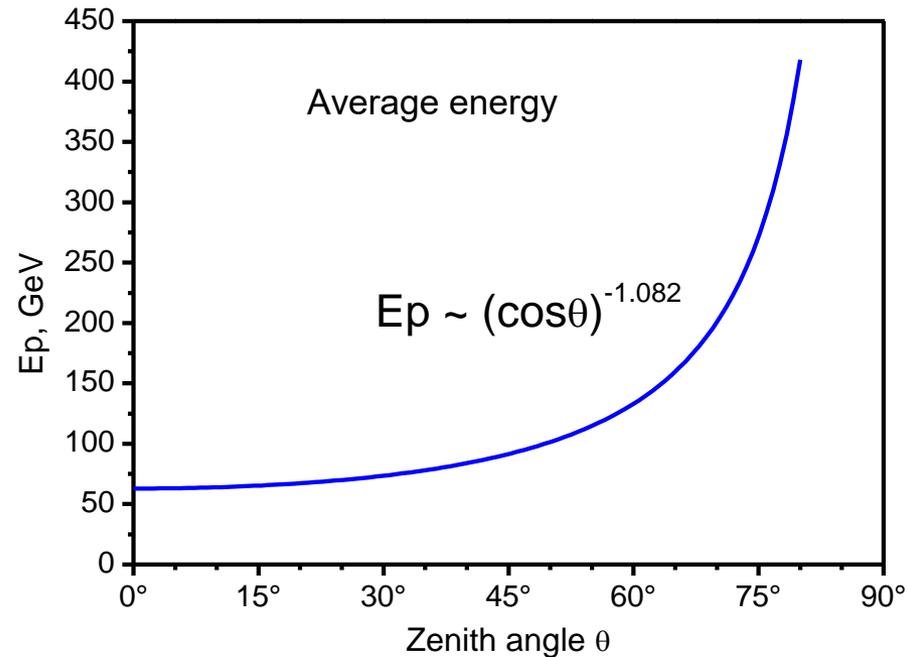
# Muons in cosmic rays

**Muons** in the atmosphere are produced mainly in decays of **pions** and **kaons**, which in their turn are produced in primary cosmic ray interactions with nuclei of atoms of the atmosphere. The ratio of energies  $E_0/E_\mu$  is about 10.



*Scheme of muon generation*

*Dependence of this energy on zenith angle*



Average muon energy at the Earth's ground

$$\langle E_\mu \rangle = 4 \text{ GeV}$$

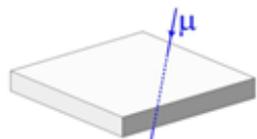
Average energy of primary protons

$$\langle E_p \rangle = 50\text{-}60 \text{ GeV}$$

Average energy of muons in near horizontal flux  $\approx 70$  GeV, and corresponding primary proton energies  $\approx 700$  GeV.

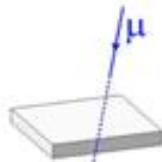
# Different types of muon detectors

Muon detector



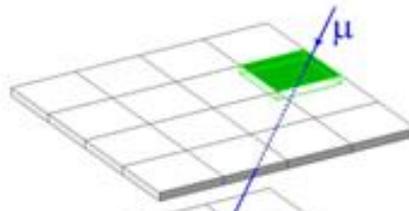
$$\Delta\theta \approx 90^\circ$$
$$\Delta\phi = 360^\circ$$

Muon telescope



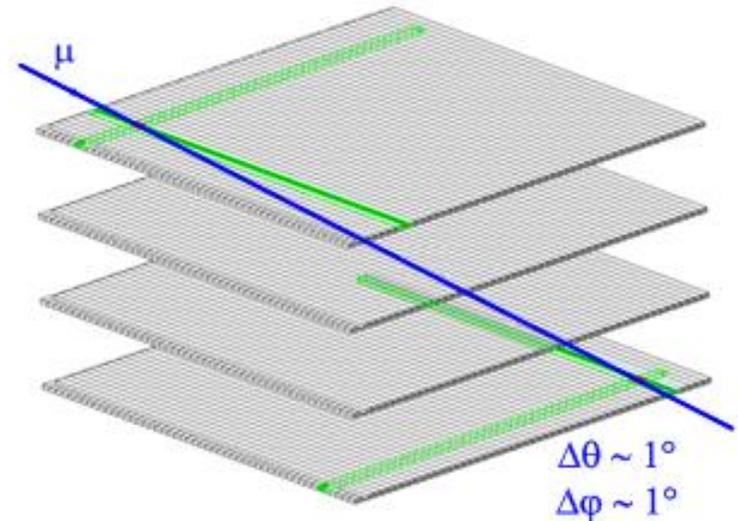
$$\Delta\theta \approx 30^\circ$$
$$\Delta\phi = 360^\circ$$

Multidirectional muon telescope



$$\Delta\theta \approx 15^\circ \div 30^\circ$$
$$\Delta\phi \approx 28^\circ \div 35^\circ$$

Muon hodoscope



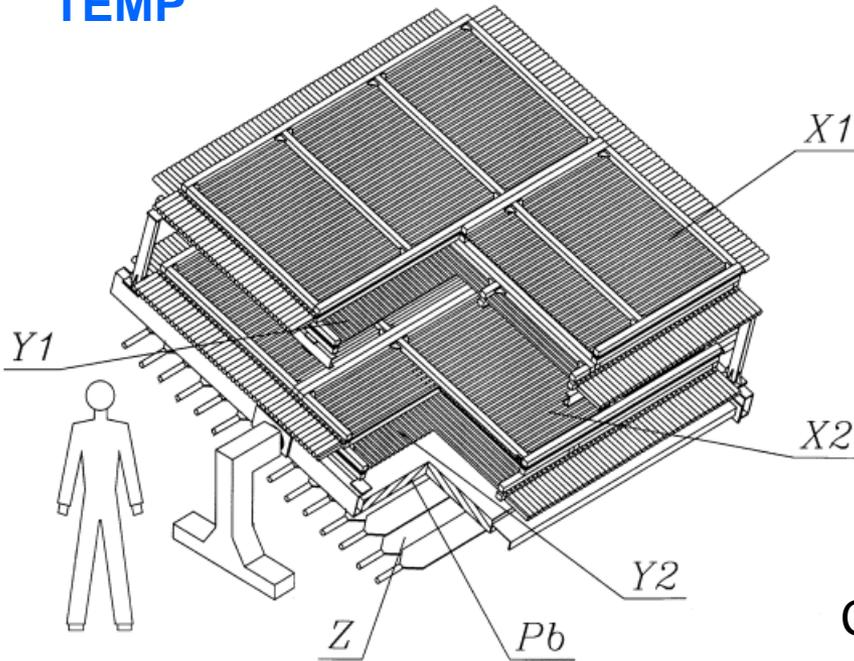
$$\Delta\theta \sim 1^\circ$$
$$\Delta\phi \sim 1^\circ$$

The main feature of muon hodoscope is track reconstruction of each muon with high spatial and angular accuracy

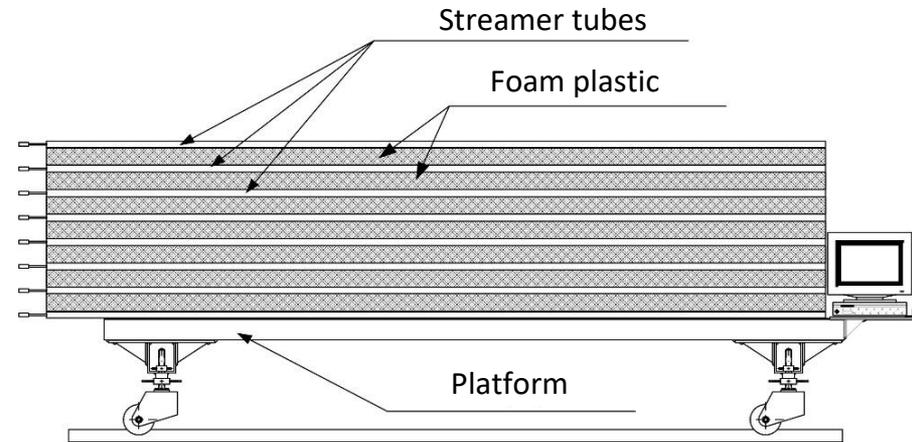
A **hodoscope** from the Greek "hodos" for way or path, and "skopos" an observer

# Muon hodoscopes for muonography

TEMP



URAGAN

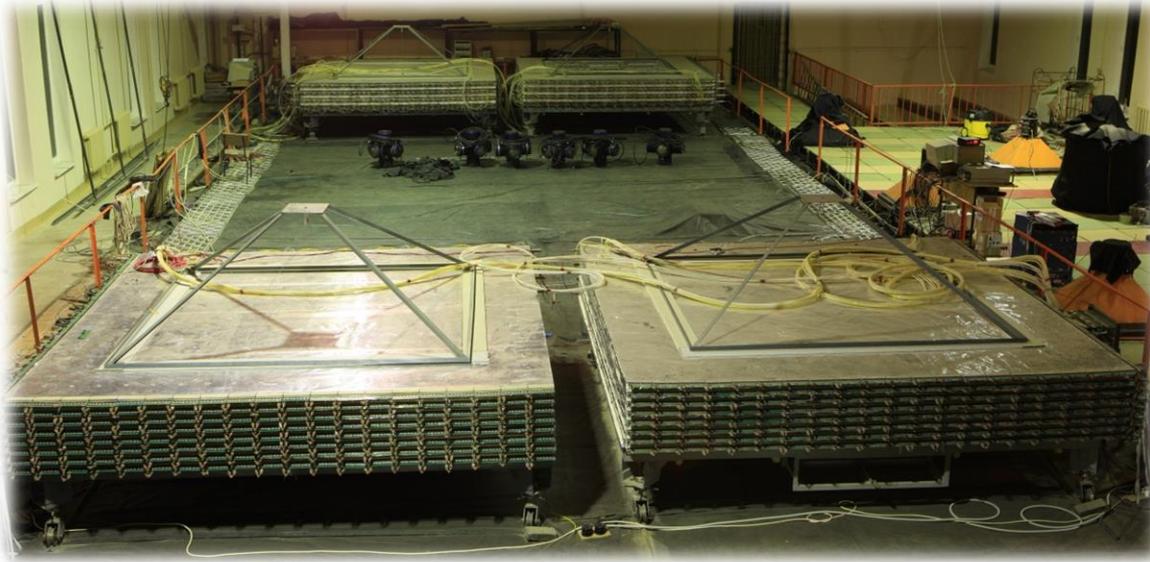


TEMP URAGAN

Coordinate planes	2	8
Area, m <sup>2</sup>	9	11.5
Angular accuracy, °	2	0.8
Counting rate, s <sup>-1</sup>	550	1400
Efficiency of registration, %	90	99

The first muon hodoscope TEMP was constructed in 1995 in MEPhI, 2 coordinate planes only . More perfect MH consisting of 8 coordinate planes was constructed also in MEPhI in 2005-2006. A similar MH also consists of 8 coordinate planes and was constructed in Mexico in 2014.

# Muon hodoscope URAGAN



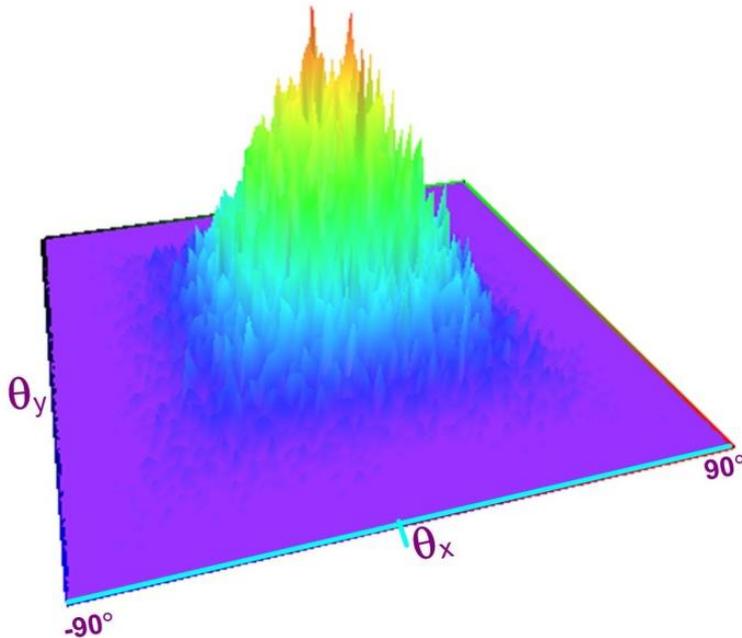
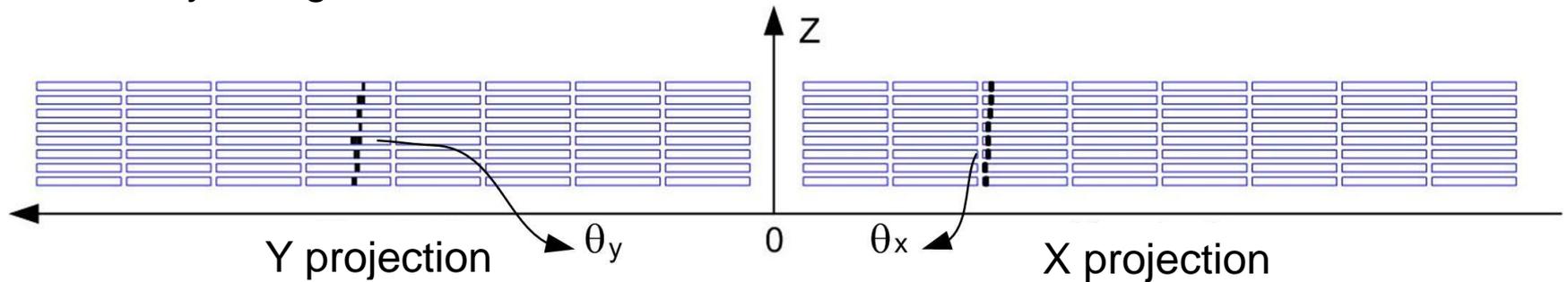
**Muon hodoscope URAGAN** consists of 4 supermodules. The total area is 45 sq. m. The supermodule provides high spatial and angular accuracy of muon track registration (1 cm and  $1^\circ$ , correspondingly) in a wide range of zenith angles: from  $0^\circ$  to  $80^\circ$ .

## Capabilities :

- **Integral mode:** measurements of the variations of total muon flux.
- **Hodoscopic mode** provides simultaneous measurements of muon flux coming from various directions of the celestial hemisphere.
- **Matrix method for storing information is used:**
  - zenith-angular dependence; azimuthal dependence; muon flux anisotropy.

# Matrix method of registration

The arrival direction of each muon is determined in real-time. Each event is added in the appropriate cell of the two-dimensional matrix of projection angles having the size of 2 by 2 degrees.

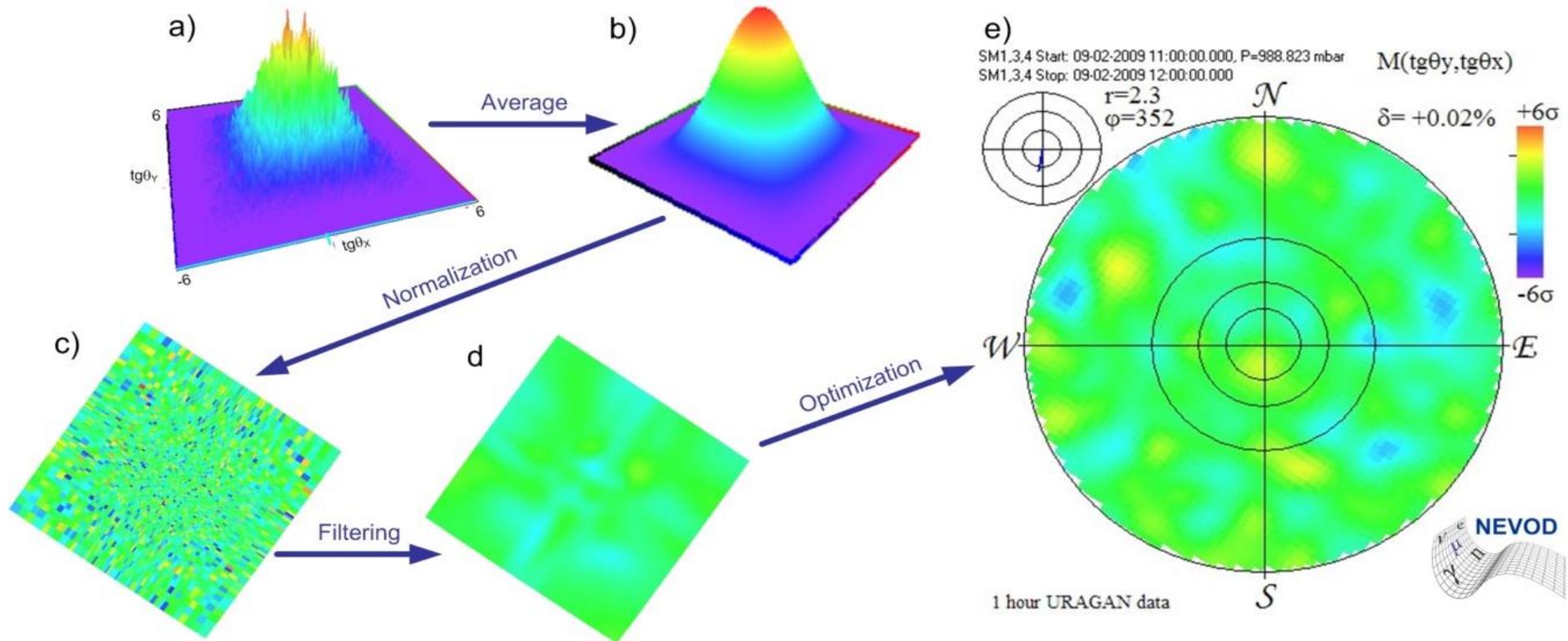


Maximum corresponds to the vertical direction.

**Every minute** one SM of the URAGAN registers about **80,000 muons**. The matrices averaged over the data of three supermodules obtained during **one hour** expositions are used. The statistical reliability of every such matrix is about **15 million events**. Statistical error is of  $\sim 0.1\%$  (for 10 minute interval).

# From muon matrix to muonography

To separate small deviation in muon flux from basic muon flux the **muonography technique** is used. To get muonographs we do the following manipulations:



*Color shows the deviation from the average in sigma (red color – excess of muons, blue color – deficit of muons)*

Such muonographs allow to study the dynamics of two-dimensional muon flux variations.

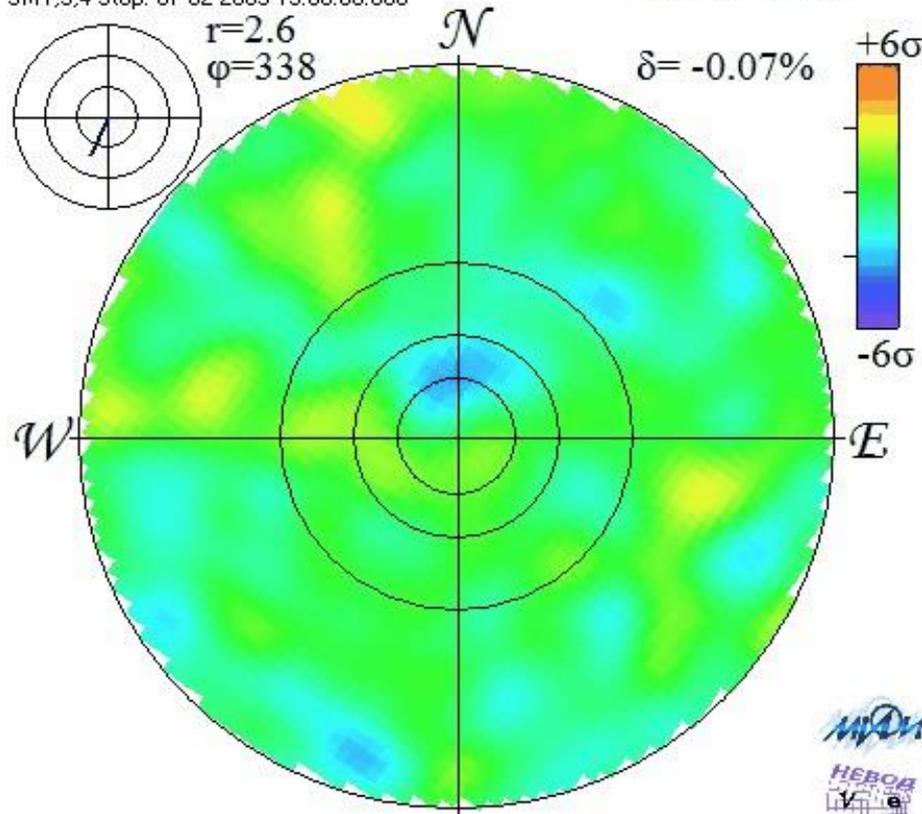
# Dynamics of muonographs in laboratory system in a quiet period

June25\_Presentation\_12\_BarbashinaN\_Anim\_1.gif

SM1,3,4 Start: 07-02-2009 14:00:00.000, P=989.053 mbar  
 SM1,3,4 Stop: 07-02-2009 15:00:00.000

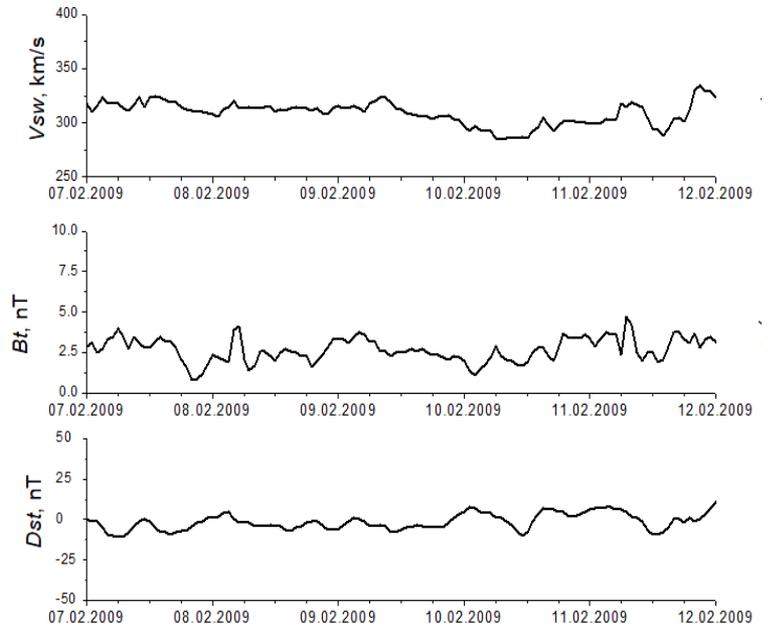
$M(\text{tg}\theta_y, \text{tg}\theta_x)$

$\delta = -0.07\%$



1 hour URAGAN corrected data  $\mathcal{S}$

No serious deviations



Parameters  
of near terrestrial space:

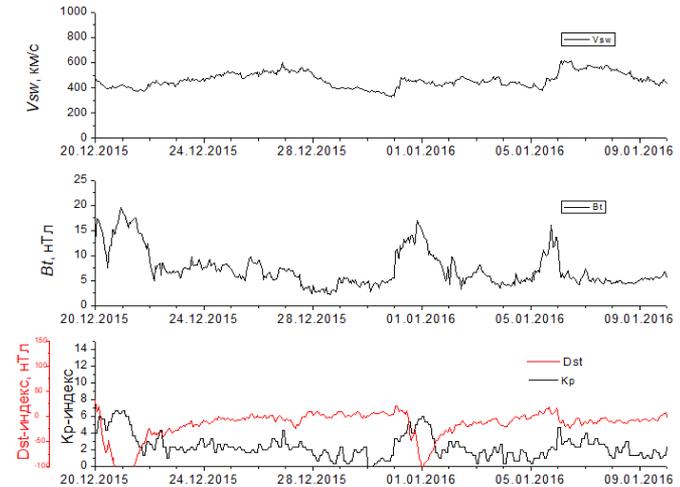
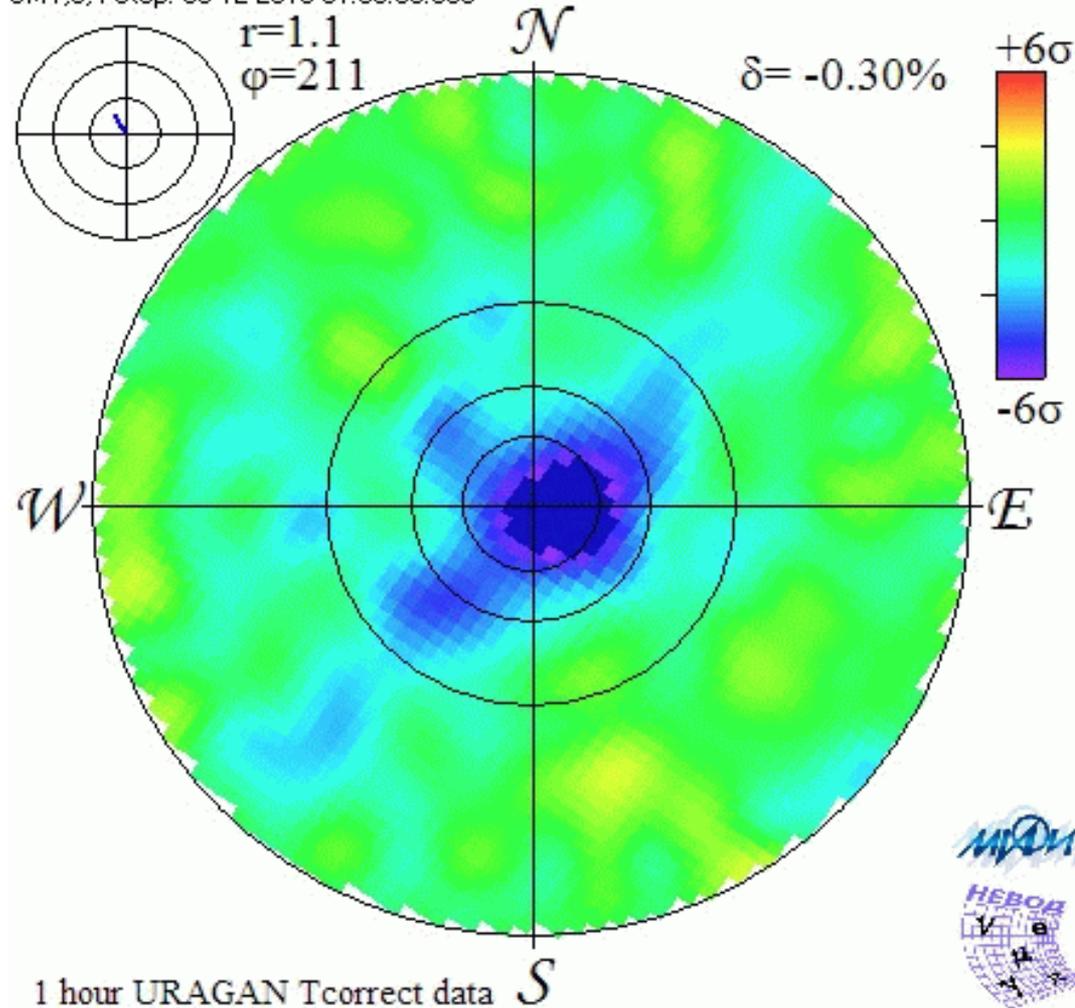
$$\langle V_{SW} \rangle \sim 325 \text{ km/s};$$

$$\langle B_{max} \rangle = 4.5 \text{ nT};$$

$$\langle Dst \rangle = 2.5 \text{ nT};$$

# Dynamics of muonographs of heliospheric disturbance in laboratory system during FD of 30 January 2015

SM1,3,4 Start: 30-12-2015 00:00:00.000, P=1016.207 mbar  
 SM1,3,4 Stop: 30-12-2015 01:00:00.000



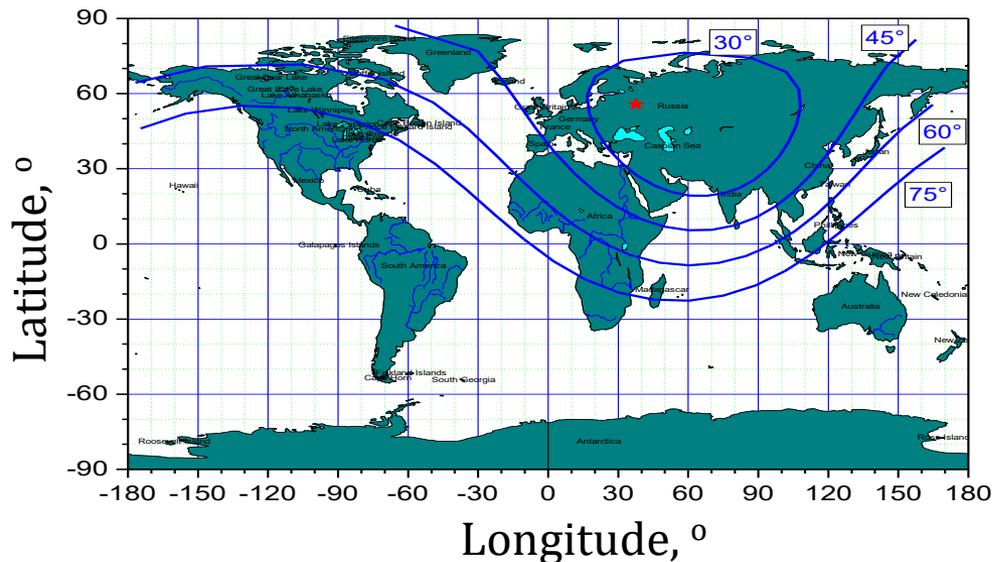
Parameters  
 of near terrestrial space:

- $\langle V_{sw} \rangle \sim 460$  km/s;
- $\langle B_{max} \rangle = 16.9$  nT;
- $\langle Dst \rangle = -100$  nT;

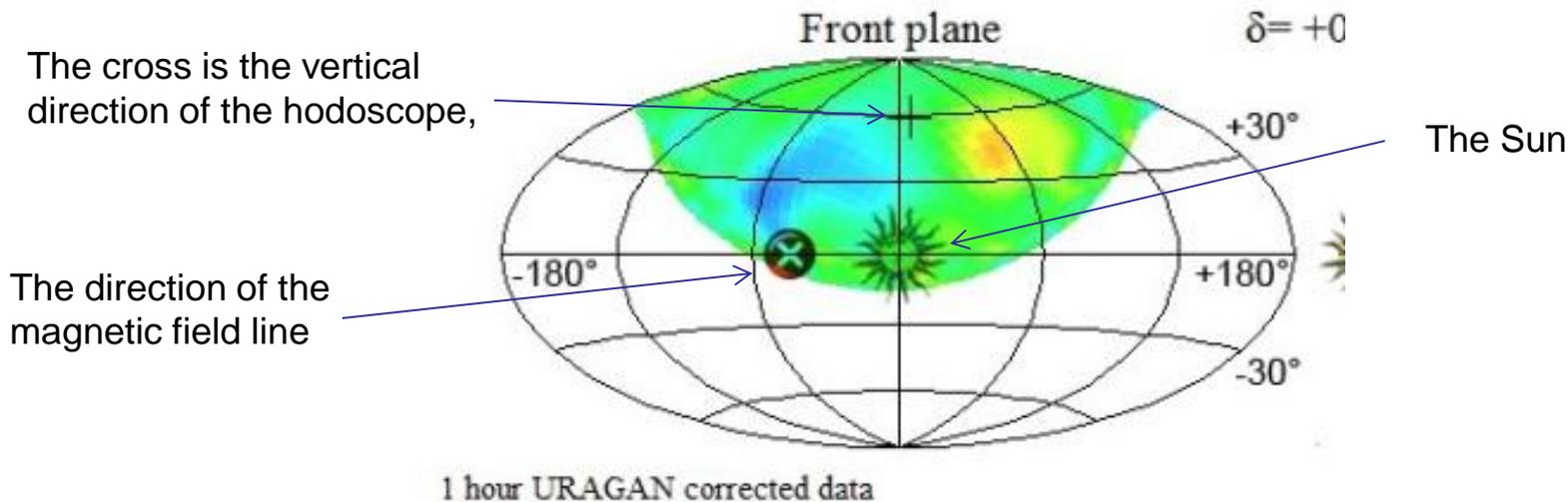
Large changes in the angular distribution of the flux of muons



# Asymptotic directions of cosmic ray protons for zenith angles of detected muons 30°, 45°, 60° и 75° by URAGAN



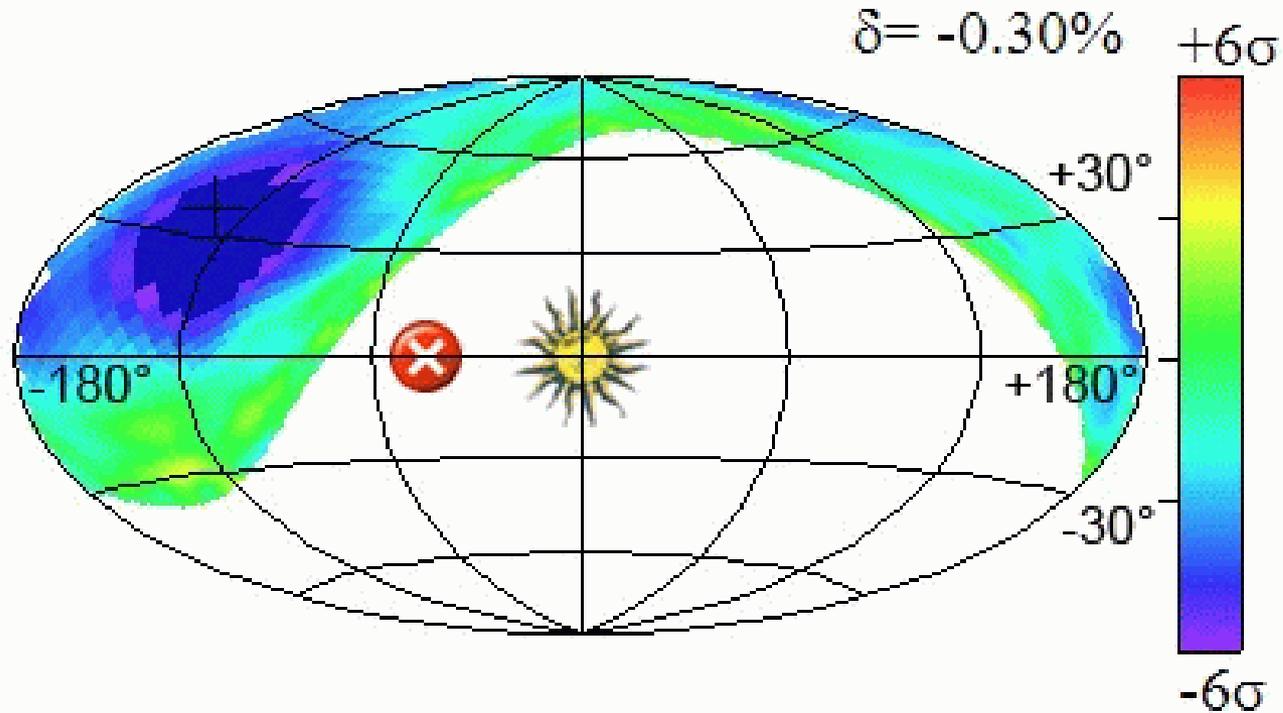
By using asymptotic directions it is possible to translate muonographs obtained in laboratory system to geocentric solar ecliptic system



# Dynamics of muonographs of heliospheric disturbance in GSE system during FD of 30 January 2015

SM1,3,4 Start: 30-12-2015 00:00:00.000, P=1016.207 mbar  
SM1,3,4 Stop: 30-12-2015 01:00:00.000

$M(\text{tg}\theta_y, \text{tg}\theta_x)$



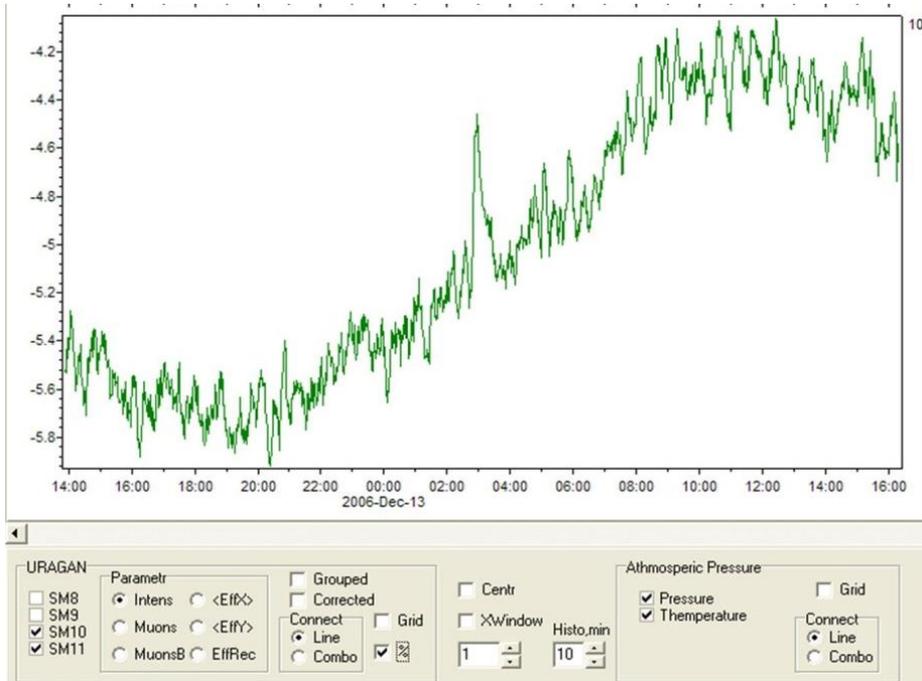
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1 hour URAGAN Tcorrect data

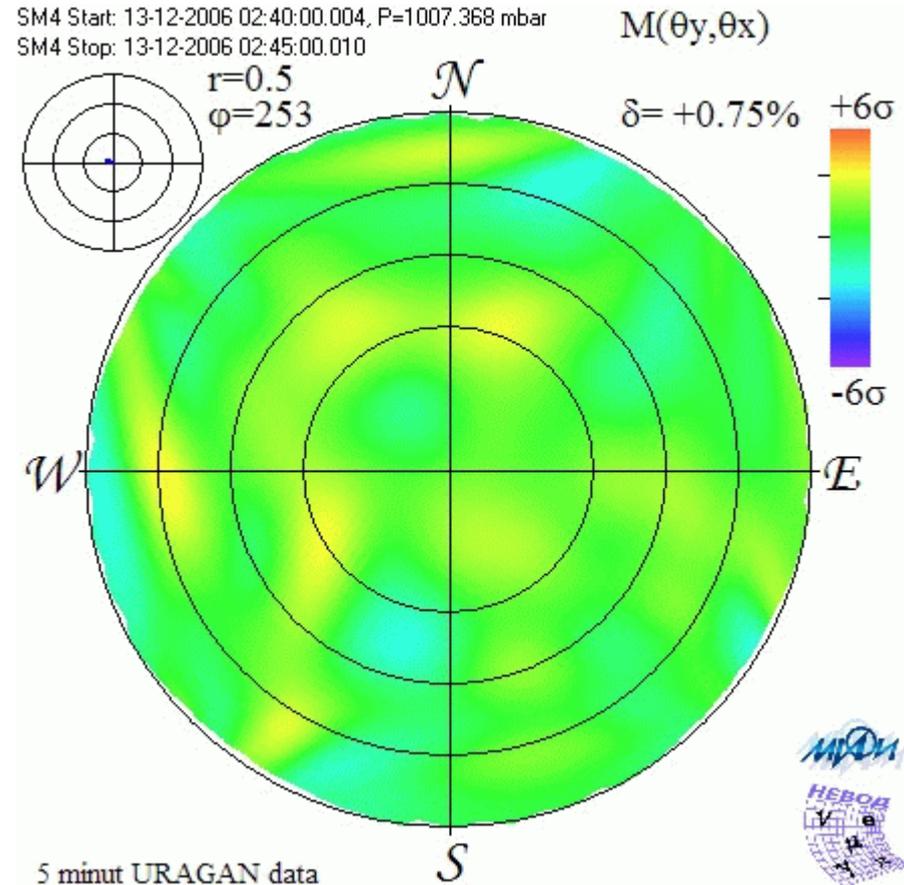


# GLE #70 of December 13, 2006

This was very interesting event and muonographs allow to see angular distribution during the increase of muon flux in laboratory system.



*The integral counting rate gives only the time and amplitude of the increase*



*Muonographs allow to see the directions of these changes*

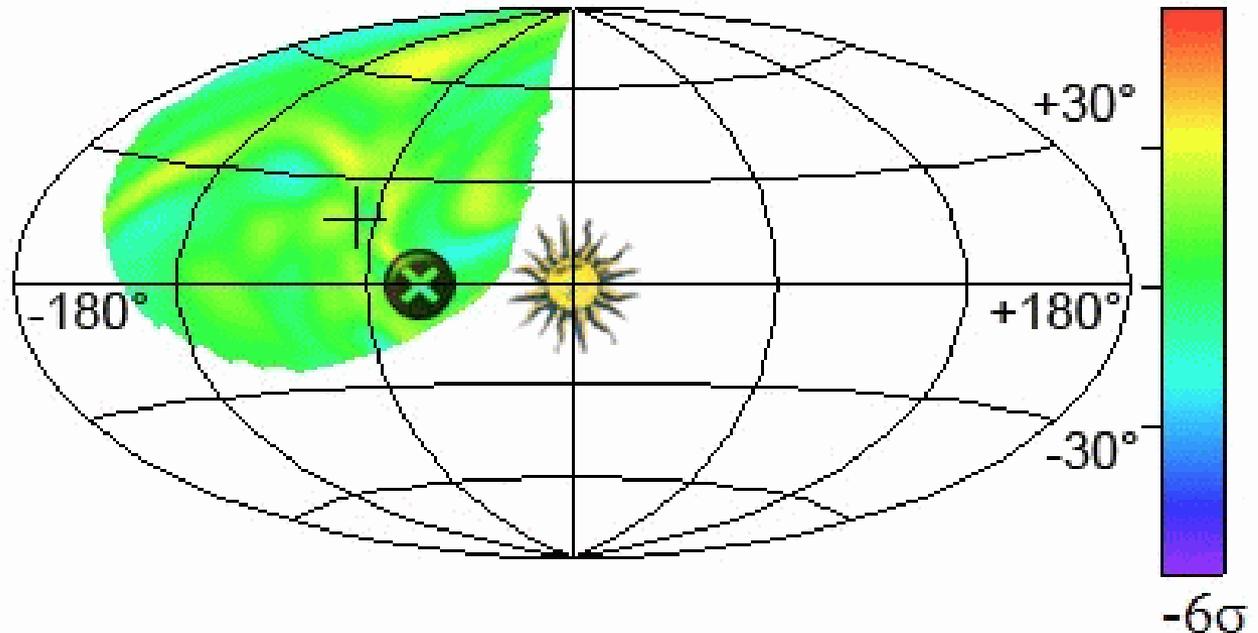
# GLE #70 of December 13, 2006 in GSE

SM4 Start: 13-12-2006 02:40:00.004, P=1007.368 mbar

SM4 Stop: 13-12-2006 02:45:00.010

$M(\theta_y, \theta_x)$

$\delta = +0.75\%$   $+6\sigma$



[June25\\_Presentation\\_12\\_BarbashinaN\\_Anim\\_5.gif](#)

5 minut URAGAN data

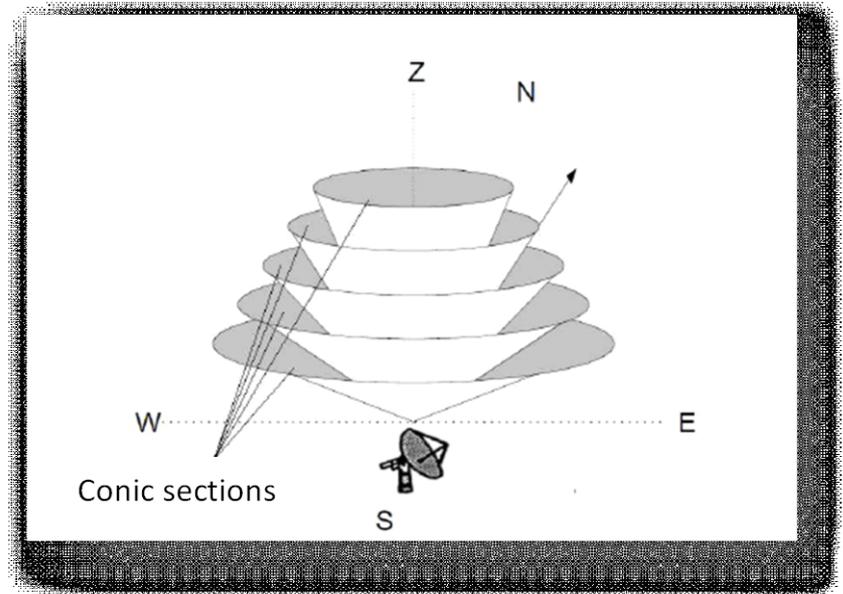


Maximum muon flux is seen in the direction near to magnetic field line

# Muonography of atmospheric processes

Thunderstorms are a clear manifestation of non-stationary processes in the atmosphere. For calibration of muonography method, independent information about thunderstorms from Doppler weather radar DMRL-C is used.

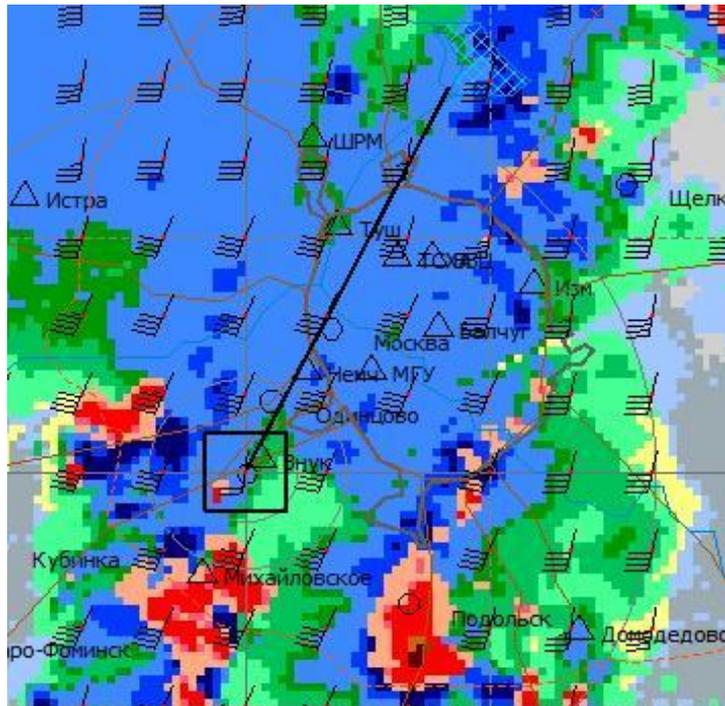
## Doppler weather radar DMRL-C



Radio locator determines the radio beam reflectivity from various hydrometeors (droplets, snowflakes, etc.). Instrumental range of radio beam is 250 km, maximum detection height is 20 km.

## DMRL-C data

Radio locator data represent 3-dimensional maps of radar scanning (map of phenomena). Resolution of maps is  $1 \times 1 \text{ km}^2$ .



*Example of thunderstorm event on July 27, 2015 on the map of meteorocator. Black square represents the locator position, inclined vertical line points the direction of air mass movement. Red color corresponds thunderstorm.*

# Comparison of Doppler maps and muonographs

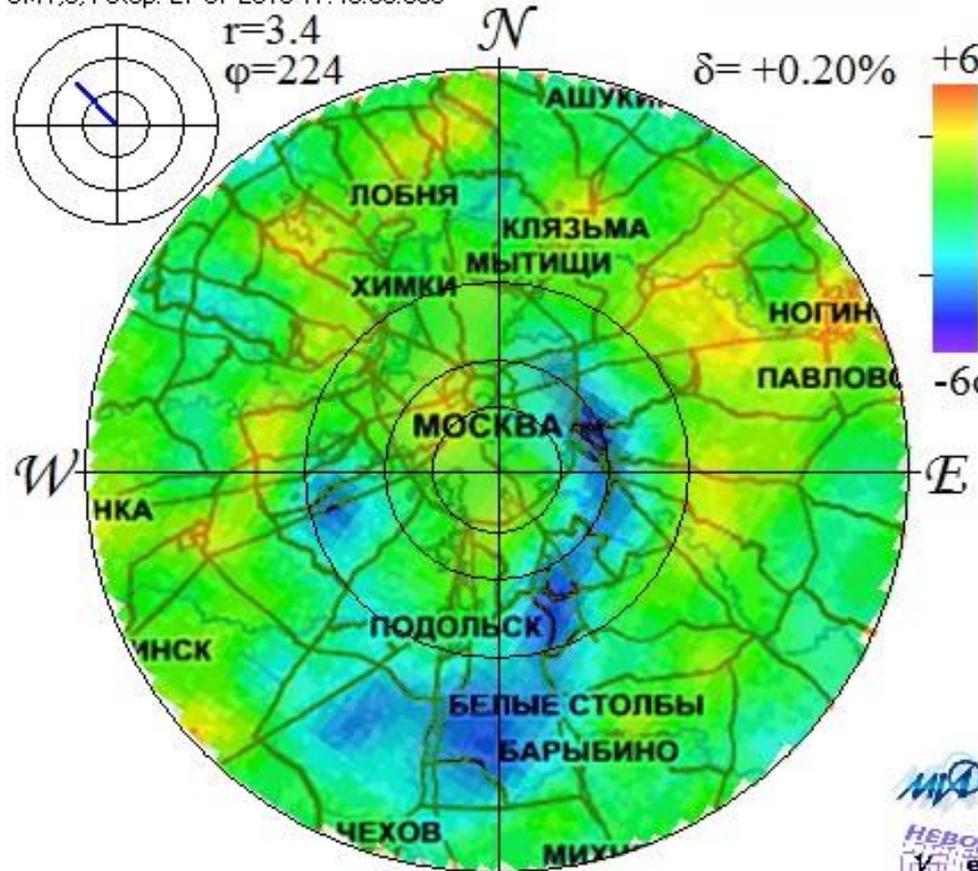
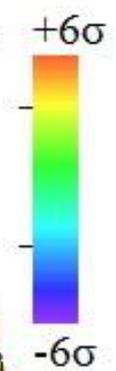
SM1.3.4 Start: 27-07-2015 17:40:00.000, P=987.602 mbar

SM1.3.4 Stop: 27-07-2015 17:45:00.000

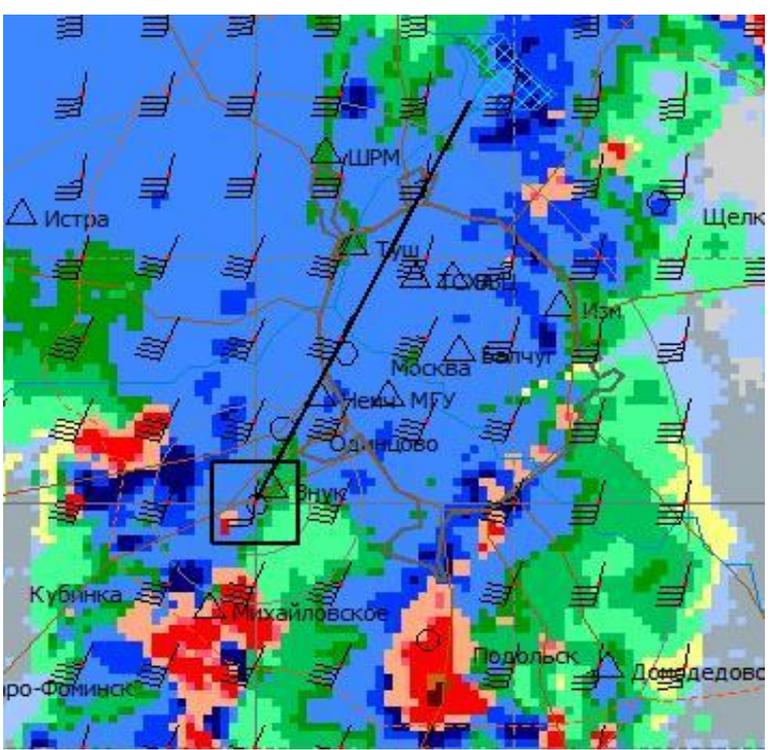
$M(\text{tg}\theta_y, \text{tg}\theta_x)$

$r=3.4$   
 $\varphi=224$

$\delta = +0.20\%$

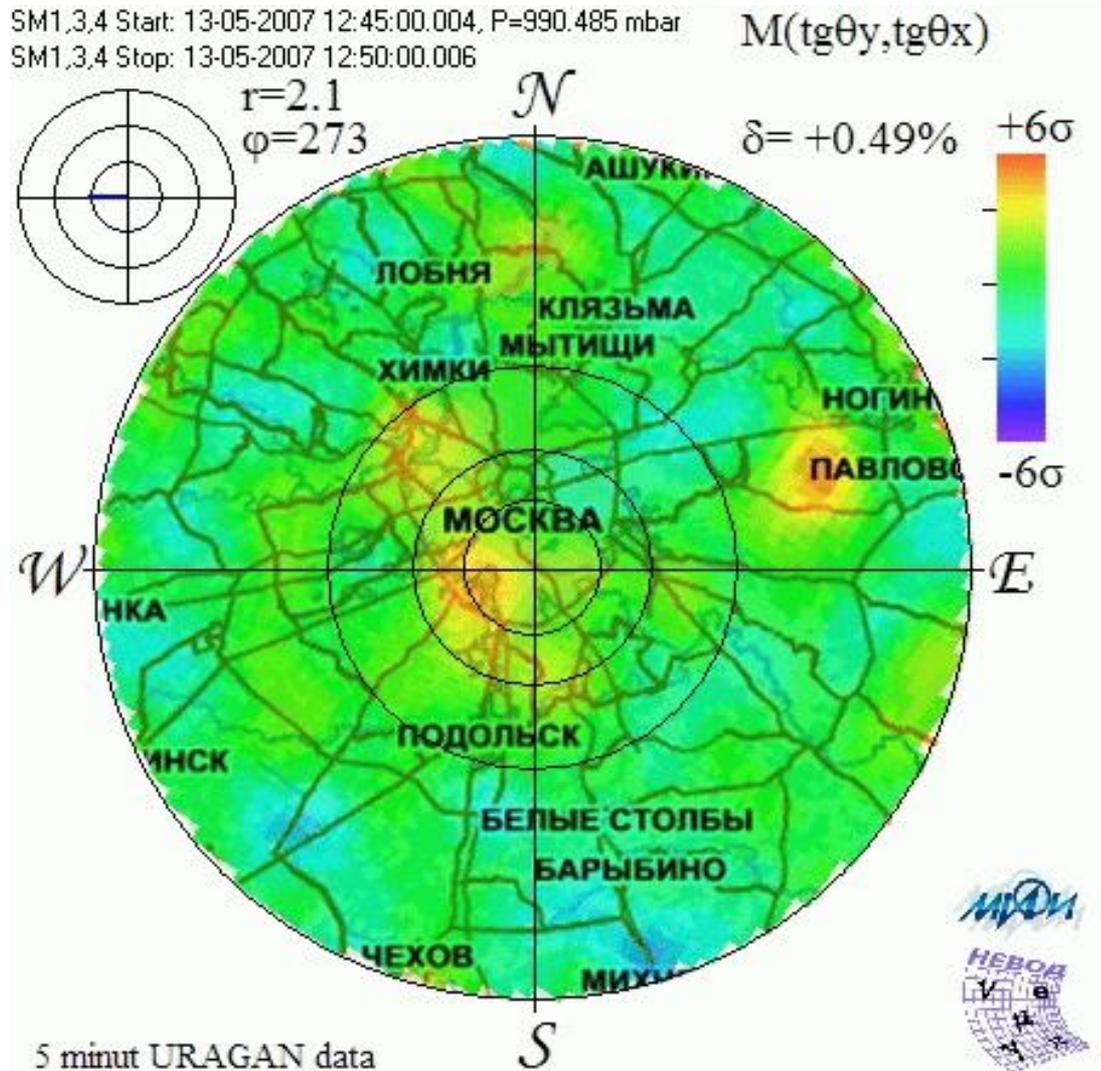


5 minut URAGAN corrected data



This example shows that the area of thunderstorm corresponds to minimum muon flux.

# Dynamics of muonographs of atmospheric disturbances



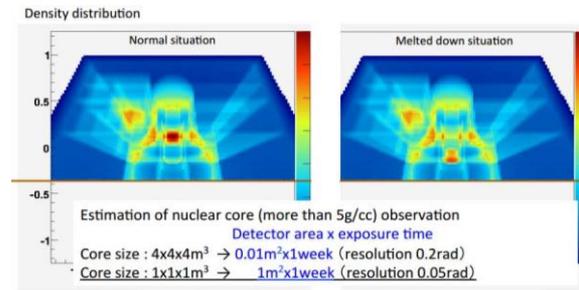
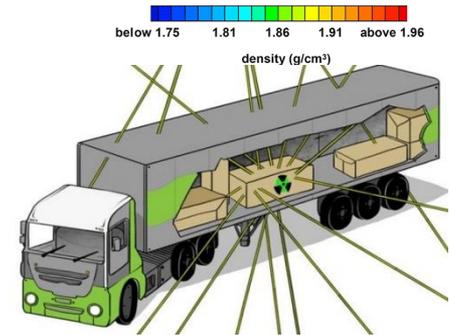
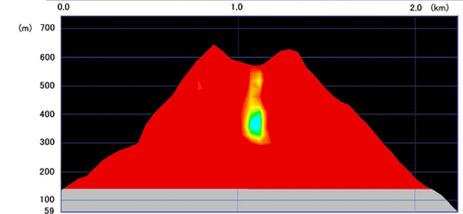
This example shows changes of muonographs during thunderstorm in time.

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For more details see talk Kachur Alexandra on Friday, 28 June, Session 5B

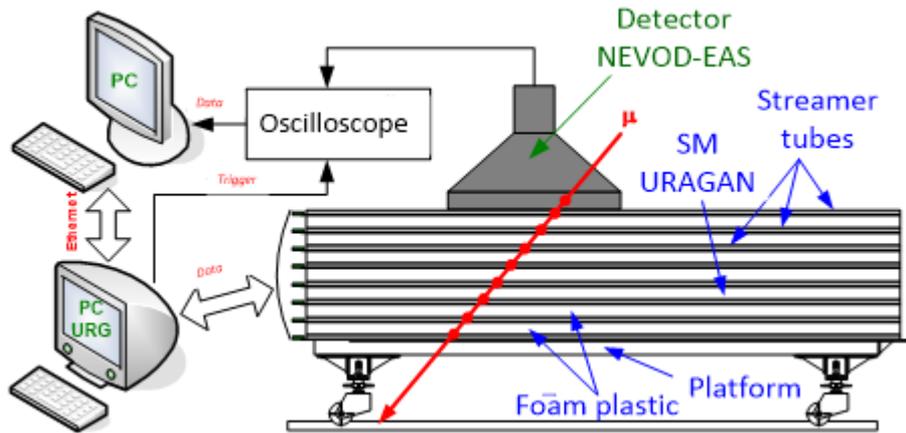
# Muonography of Earth's objects (muon tomography)

- Study of the structure of Egyptian (Cheops, Chephren, etc.) and Mexican (Teotihuacan) pyramids.
- Mapping of the internal structure of volcanoes (Vesuvius, Stromboli, Etna and Puy de Dome, Satsuma-Iwojima) for the magma level estimations.
- Prevention of forbidden transportation of heavy materials.
- Muonography of the location of nuclear fuel in emergency reactors of the NPP (Fukushima-1).
- Muonography of particle detectors.



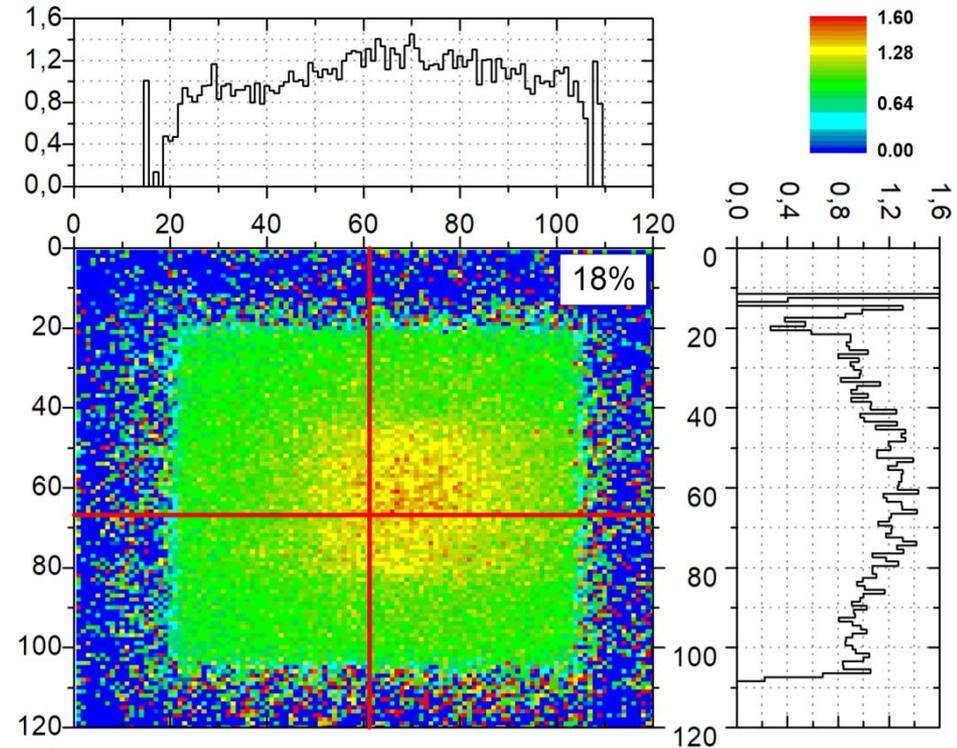
# The muonography of particle detectors (non-uniformity of the scintillation counter light collection)

Scheme of measurements



Color corresponds to **relative counter response** to the passage of single muon through the cell with dimensions of  $1 \times 1 \text{ cm}^2$ .

The muonography of the counter response



**Yellow region** corresponds to the passage of muons through the PMT

# Conclusion

1. Muon hodoscope technique and method of muonography allow to implement a new approach to investigations of heliospheric, magnetospheric and atmospheric disturbances, and to study the structure of various objects.
2. Results of observations with the help of muonography in the real-time mode can be found at:
  - in the atmosphere  
<http://nevod.mephi.ru/English/sample2.php>
  - and in the heliosphere  
<http://nevod.mephi.ru/English/sample3.php>

**Thank you very much!**